## Feedback of zonal flows on Rossby-wave turbulence driven by small scale instability

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## Coexistence of large scale coherent structures and small scale turbulence in geophysical flows





Zonal turbulence on Jupiter (NASA)

Eddy-resolving simulation of Earths oceans (Earth Simulator Center/JAMSTEC).

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## Zonal Flows and turbulence in magnetised plasmas





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Plasma turbulence (L. Villard)

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Introduction: Rossby wave turbulence

#### 2 Generation of large scales by modulational instability

#### Feedback of large scales on small scales







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## Charney-Hasegawa-Mima equation driven by small scale "instability"

$$(\partial_t - \mathcal{L}) \left( \Delta \psi - F \psi \right) + \beta \, \partial_x \psi + J[\psi, \Delta \psi] = \mathbf{0}.$$

Operator  $\mathcal{L}$  has Fourier-space representation

$$\mathcal{L}_{\mathbf{k}} = \gamma_{\mathbf{k}} - \nu k^{2m}$$

which mimics a small scale instability (and hyperviscosity).



- $\gamma_{\mathbf{k}}$  peaked around  $(k_f, 0)$ .
- Choose  $k_f^2 \sim F$ .
- Forcing excites meridionally propagating Rossby waves with wavelength comparable to deformation radius.

### Large scale – small scale feedback loop

Initialise vorticity field with white noise of very low amplitude. What are the dynamics? The following scenario was proposed by Zakharov and coworkers (1990's):

- Waves having  $\gamma_{\mathbf{k}} > 0$  initially grow exponentially.
- When amplitudes get large enough, nonlinearity initiates cascades.
- Inverse cascade transfers energy to large scales leading to zonal jets.

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 Jets shear small scale waves generating negative feedback. "Switches off" the forcing.

This talk: is this what really happens?

## Weak and strong turbulence regimes

Nonlinearity measured by dimensionless amplitude:

$$M = \frac{\Psi_0 k^3}{\beta}$$

where k typical scale,  $\Psi_0$  is typical amplitude. Two limits:

- $M \gg 1$  : Euler limit.
- $M \ll 1$ : Wave turbulence limit.

Wave turbulence limit is analytically tractable:

- Evolution of turbulence spectrum described by a closed kinetic equation.
- Kinetic equation has exact stationary solutions describing cascades.

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## Wave turbulence kinetic equation

$$\frac{\partial n_{\mathbf{k}}}{\partial t} = 4\pi \int \left| V_{\mathbf{q}\,\mathbf{r}}^{\mathbf{k}} \right|^{2} \delta(\mathbf{k} - \mathbf{q} - \mathbf{r}) \delta(\omega_{\mathbf{k}} - \omega_{\mathbf{q}} - \omega_{\mathbf{r}}) \times [n_{\mathbf{q}}n_{\mathbf{r}} - n_{\mathbf{k}}n_{\mathbf{q}}\operatorname{sgn}(\omega_{\mathbf{k}}\omega_{\mathbf{r}}) - n_{\mathbf{k}}n_{\mathbf{r}}\operatorname{sgn}(\omega_{\mathbf{k}}\omega_{\mathbf{q}})] d\mathbf{q}d\mathbf{r} + \gamma_{\mathbf{k}}n_{\mathbf{k}}.$$

where

$$\omega_{\mathbf{k}} = -\frac{\beta \, k_x}{k^2 + F}$$
$$V_{\mathbf{q}\,\mathbf{r}}^{\mathbf{k}} = \frac{i}{2} \sqrt{\beta \, |k_x q_x r_x|} \left( \frac{q_y}{q^2 + F} + \frac{r_y}{r^2 + F} - \frac{k_y}{k^2 + F} \right).$$

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**Problem:** stationary Kolmogorov solution describing the inverse cascade is non-local.





#### 2 Generation of large scales by modulational instability

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## Modulational instability of Rossby waves

Small scale Rossby waves are unstable to large scale modulations (Lorenz 1972, Gill 1973)



 In wave turbulence limit, *M* ≪ 1, unstable perturbations concentrate on resonant manifold:

> $\mathbf{p} = \mathbf{q} + \mathbf{r}$  $\omega(\mathbf{p}) = \omega(\mathbf{q}) + \omega(\mathbf{p}_{-}).$

• Perturbations with fastest growth rate become close to zonal for  $M \ll 1$ .

### Linear stage of modulational instability





Zonal velocity profile:

Growth of  $|\psi_{\mathbf{q}}|^2$  compared to predictions of linear stability.

$$U_Z(y) = \frac{1}{2\pi} \int_0^{2\pi} v_x(x, y) \, dx$$

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#### Generation of jets by modulational instability



t=12















t=18





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### Outline



#### 2 Generation of large scales by modulational instability

#### Feedback of large scales on small scales



# Distortion of small scales by large scales in weakly nonlinear regime

Modulational instability generates large scales directly from small scale waves (non-locality).

Subsequent evolution of the small scales can be described by a nonlocal turbulence theory:

- Assume major main contribution to collision integral for small scales comes from modes **q** having *q* ≪ *k* (scale separation).
- Taylor expand in **q**. Leading order equation for small scales is an anisotropic diffusion equation in **k**-space:

$$\frac{\partial n_{\mathbf{k}}}{\partial t} = \frac{\partial}{\partial k_i} S_{ij}(k_1, k_2) \frac{\partial n_{\mathbf{k}}}{\partial k_j}.$$

 Diffusion tensor, S<sub>ij</sub>, depends on structure of the large scales: more intense zonal flows give faster diffusion.



## More on the spectral diffusion approximation



- Diffusion tensor, *S<sub>ij</sub>* turns out to be *degenerate*
- A change of variables shows that diffusion is along 1-d curves:

$$\Omega = \frac{\beta \, k^2 \, k_x}{k^2 + F} = const$$

- *Wave-action*, *n*<sub>k</sub>, of a small scale wavepackets is conserved by this motion.
- Energy, ω<sub>k</sub> n<sub>k</sub>, of small scale wavepackets is not conserved.

#### Numerical view of spectral transport



Wave-action diffuses along the  $\Omega$  curves from the forcing region until it is dissipated at large wave-numbers.

## Mechanism for saturation of the large scales

- Energy lost by the small scales is transferred to the large scale zonal flows which grow more intense.
- More intense large scales results in a larger diffusion coefficient.
- Increases the rate of dissipation of the small scale wave-action → negative feedback.

#### Prediction:

Growth of the large scales should suppress the small scale instability which turns off the energy source for the large scales leading to a saturation of the large scales *without large scale dissipation*.

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## Numerical observations of feedback loop



- Feedback loop is very evident in weakly nonlinear regime.
- Scaling arguments allow one to predict the saturation level of the large scales in terms of the small scale growth rate:

 $[U(\nabla U)]_{\rm LS} \sim \gamma \beta \quad M \ll 1$ 

 $[U(\nabla U)]_{\rm LS} \sim \gamma U \quad M \gg 1$ 

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These estimates are reasonably supported by numerics especially the cross-over from the weak to strong turbulence regimes (see inset).

## Conclusions

- The CHM equation driven by "instability" forcing behaves differently than one might expect.
- There is no inverse cascade but rather direct generation of large scales by modulational instability (this could change with broadband forcing).
- Evolution of subsequent turbulence is non-local in scale and partially tractable analytically in the wave turbulence limit.
- A negative feedback loop is set up whereby the growth of the small scales is suppressed by the large scales leading to a saturation of the large scales without any large scale dissipation mechanism.

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